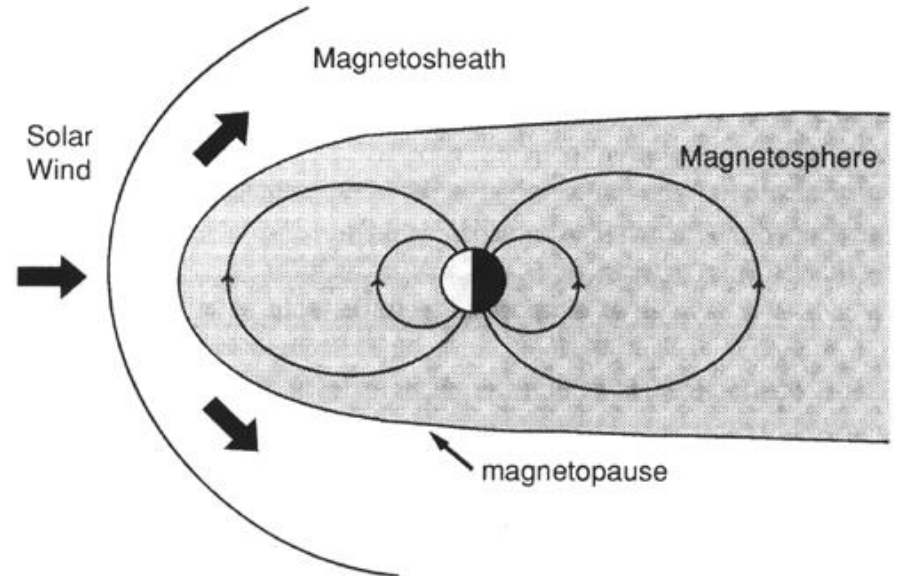
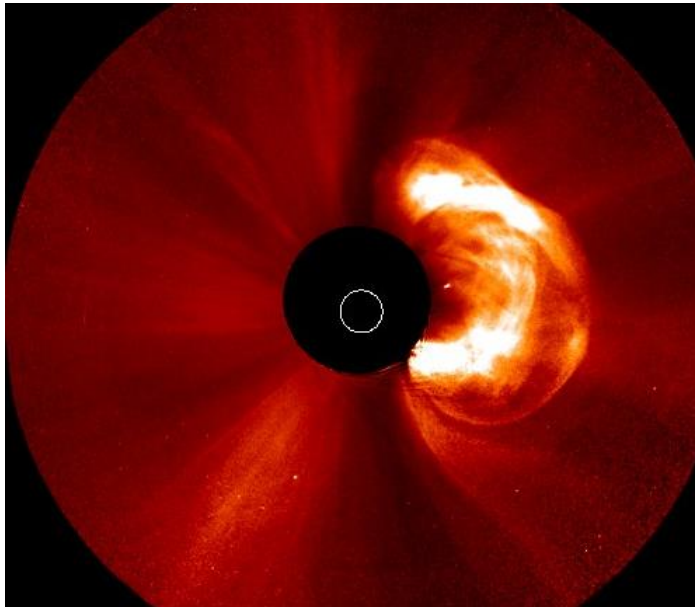


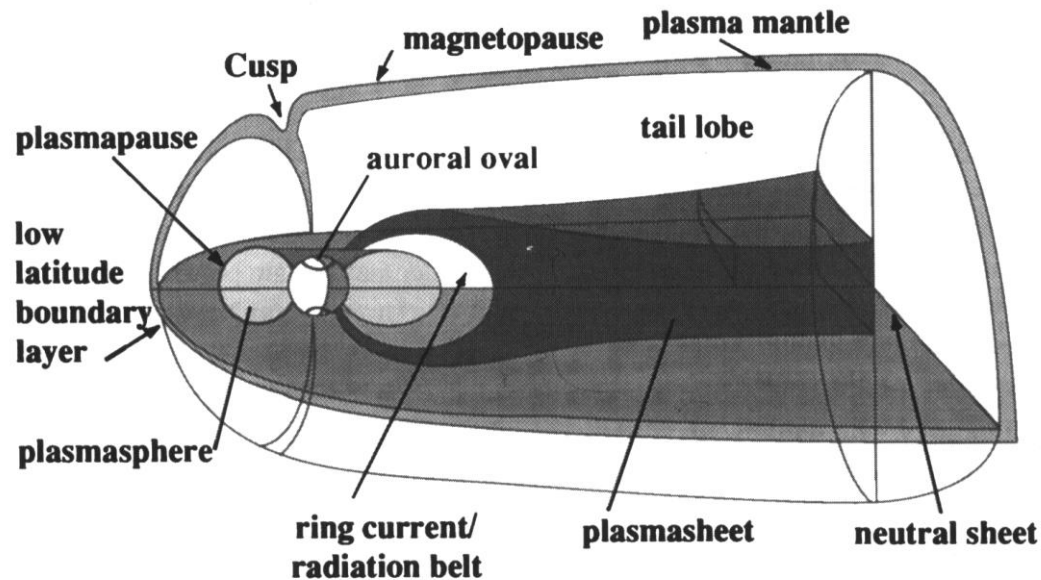
# Some Introductory Remarks on Space Weather

Iver Cairns (University of Sydney)



# 1. Space Weather

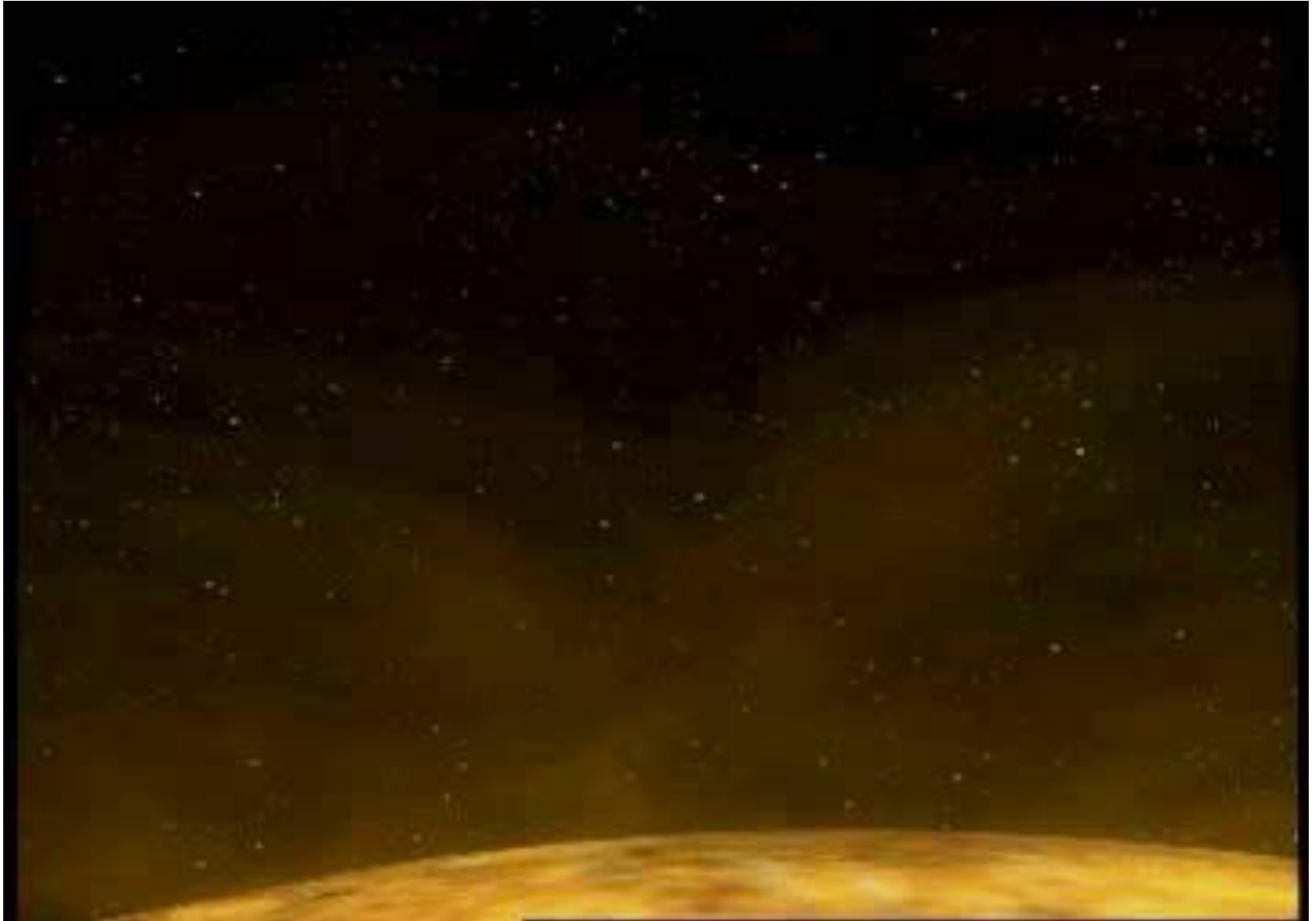
- Conditions in space that may affect human systems & activities on Earth & in space.



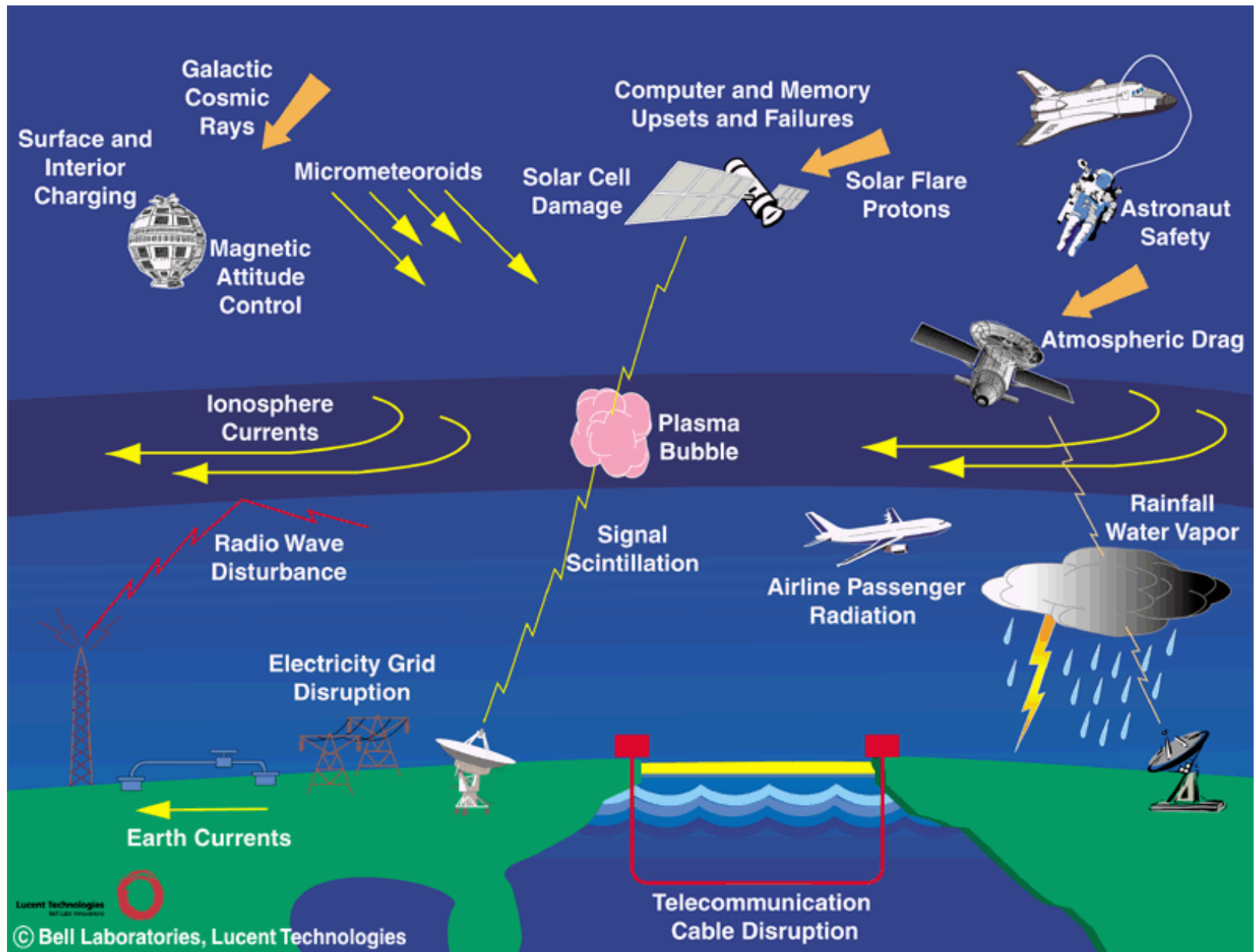
## 2. Qualitative Overview of Space Weather

- Refers broadly to conditions in space that may affect human activities on Earth & in space
- These conditions change in response to
  - Solar activity like flares (UV and X-ray)
  - Solar wind phenomena like southwards  $B_z$ ,  $P_{\text{ram}}$ , CMEs, & CIRs
  - Changes in coupling between ionosphere, magnetosphere & solar wind.
- The magnetosphere is a dynamic, time-varying system which responds in complex, not always understood ways to changes in the solar wind.

Space weather at Earth mostly due to CMEs & southward Bz  
..... due to ``magnetic reconnection’’



# Overview of Space Weather at Earth



# Terrestrial space weather events due to

- **Changes in the ionosphere** assoc. with
  - Auroral activity
  - Particle precipitation
  - Different coupling to the magnetosphere
  - Changing radio propagation conditions
- **Enhanced transport into magnetosphere** of
  - Solar wind plasma
  - Energetic particles
  - Reconnected magnetic field lines
- **Magnetic substorms and assoc. changes in **B**, **J**, & **E** in the**
  - Magnetosphere
  - Ionosphere
  - At Earth's surface

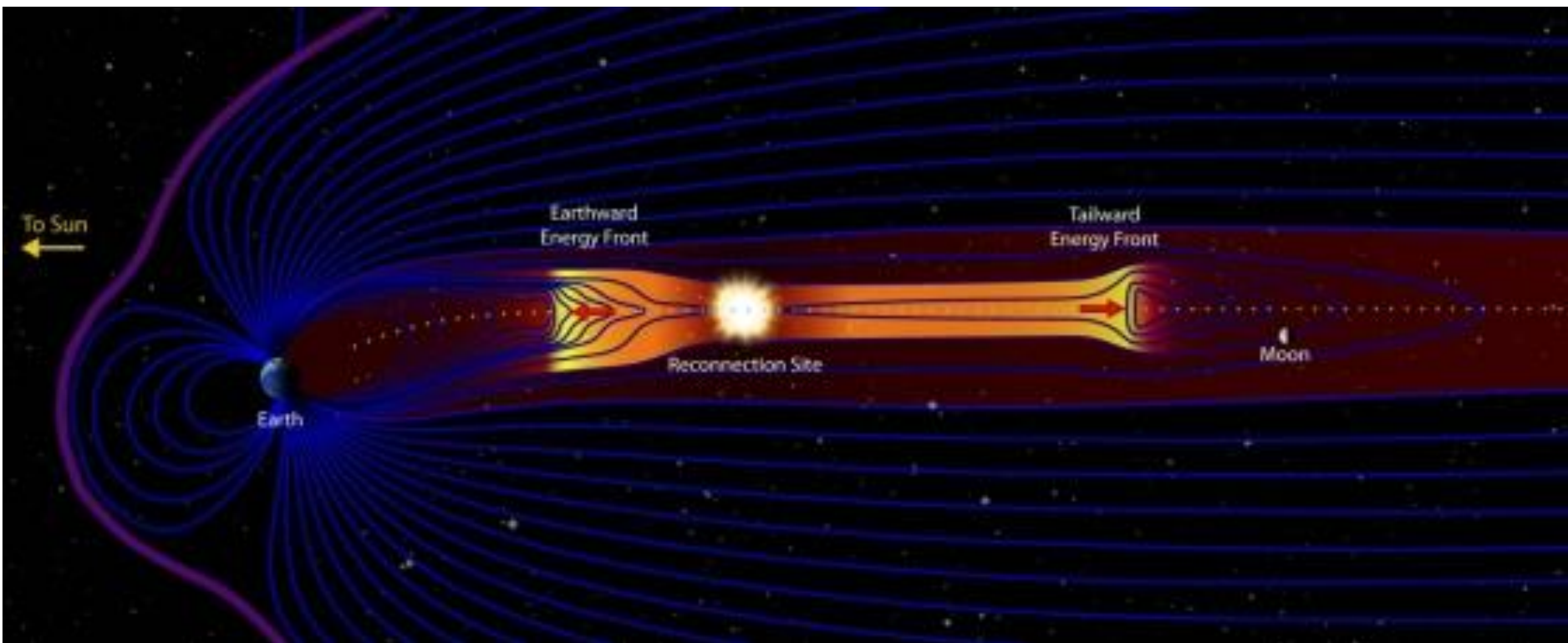
# Causes for particular human concern

- Radio, GPS, and navigation difficulties due to ionospheric scintillations & changed radio propagation
- Magnetic field changes → induced EMFs and currents → power, communication, & pipeline failures and difficulties with high-tech industry & prospecting.
- Radiation damage, dielectric charging & breakdown of space systems like satellites
- Increased ionospheric drag etc. for satellites
- Radiation damage to humans (airlines, astronauts)



### 3. Magnetic Reconnection & Space Weather

- 90% of large space weather events at Earth due to CMEs with southwards  $B_z$  & associated magnetic reconnection (B energy  $\rightarrow$  heating & flows)

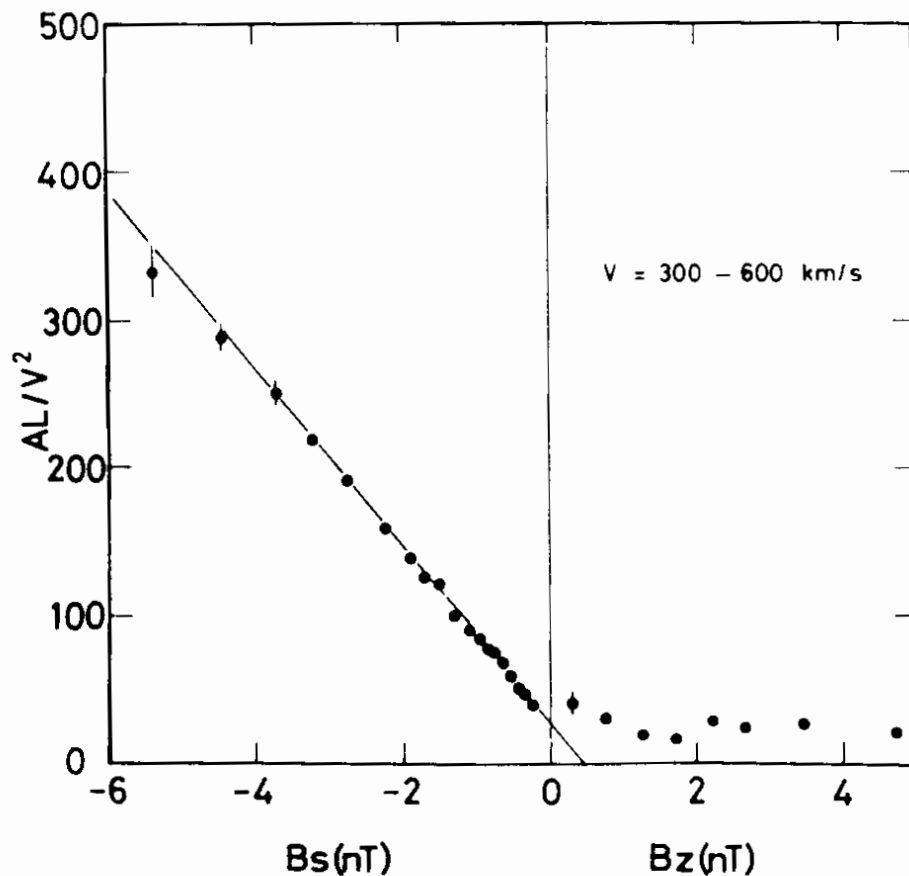




## Space weather at Earth $\Leftrightarrow$ CMEs and southwards $B_z$

- 90% of large substorms at Earth due to CMEs

[I. Richardson  
et al., 2006]



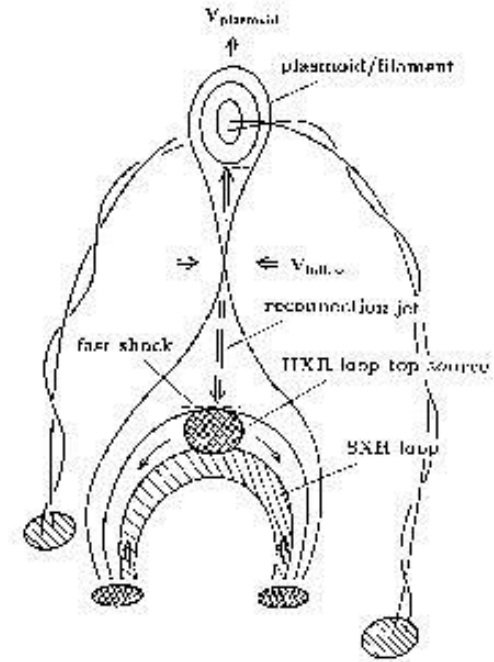
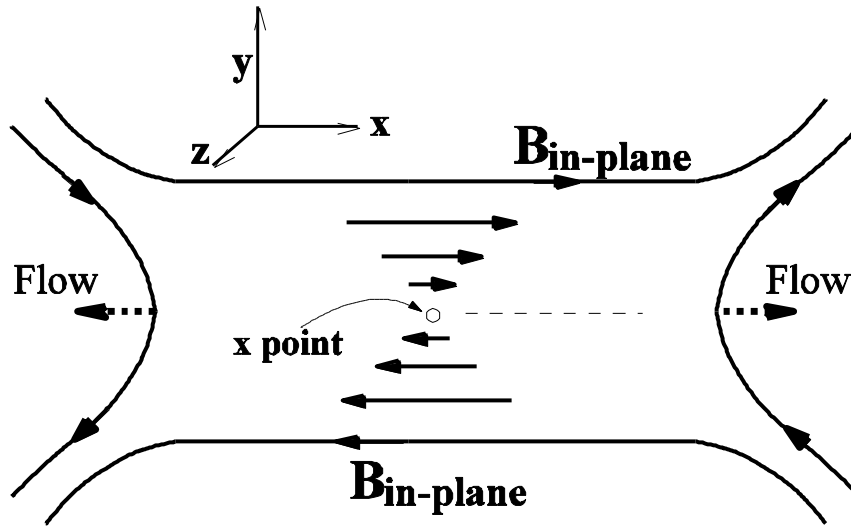
- Index  $AL / v_{sw}^2$
- $\rightarrow$  substorms assoc. with long duration and large values of southward  $B_z$
- $\rightarrow$  reconnection vital.

[Muruyama et al., 1980]

Space weather at Earth mostly due to CMEs & southward  $B_z$   
..... due to ``magnetic reconnection’’

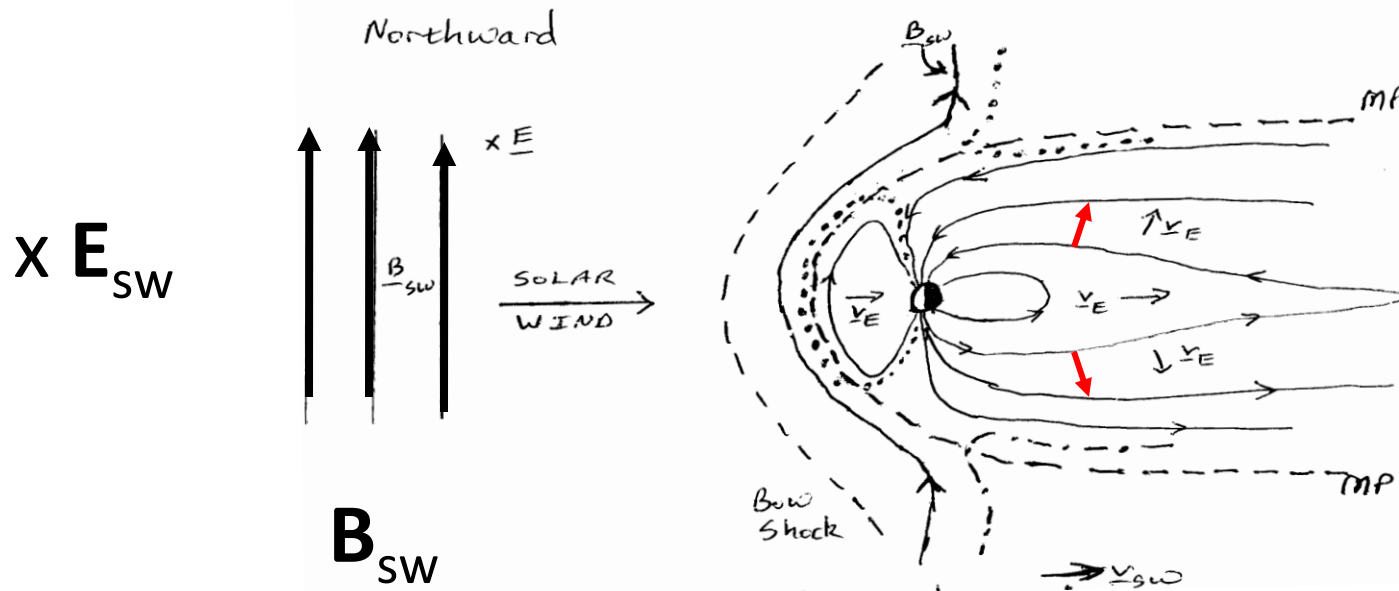


# Basics of Magnetic Reconnection



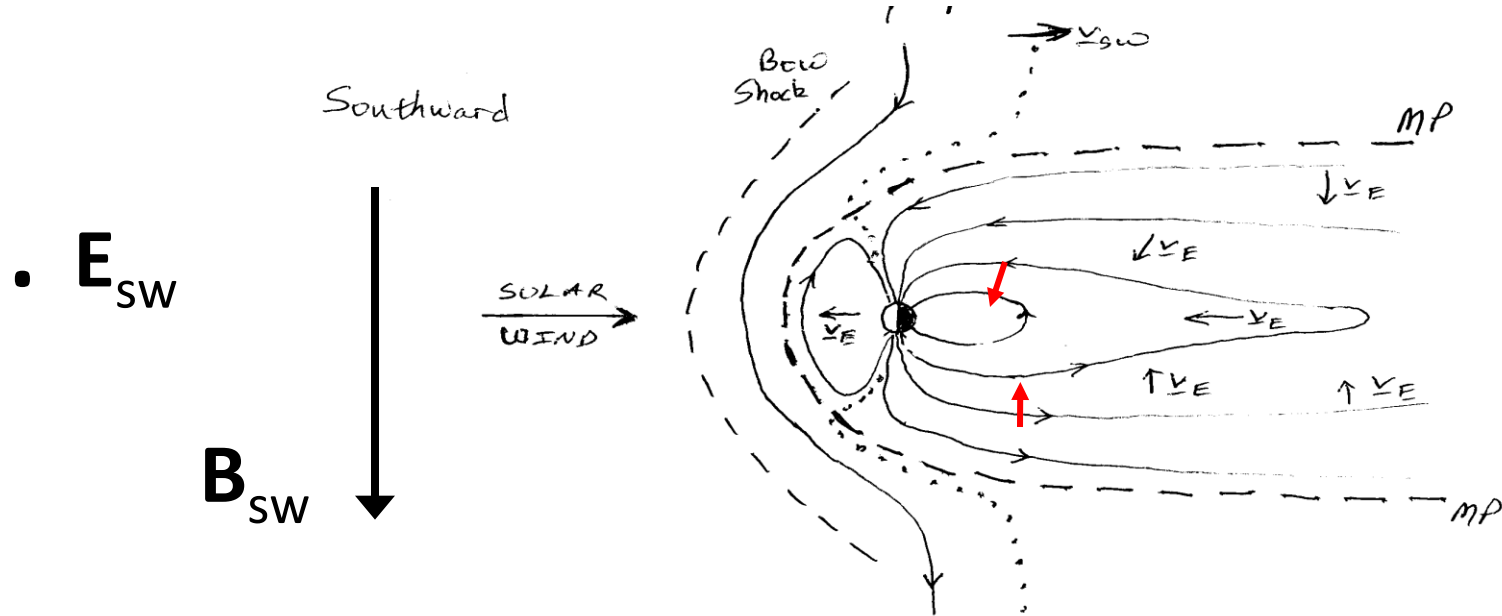
- Anti-parallel magnetic components brought together by plasma flow  $\rightarrow$   $\mathbf{B}$  annihilation.
- Where does magnetic energy  $B^2/2\mu_0$  go when  $\mathbf{B} + (-\mathbf{B}) = 0$  ?
- Some to transverse outflow at Alfvén speed  $V_A$ .
- Most  $\rightarrow$  plasma heating and energetic particle acceleration (?)

# Reconnection: Northward IMF & dusk-to-dawn $\mathbf{E}_{\text{SW}}$



- High latitude magnetic reconnection only
- Load  $\mathbf{B}$  into dayside & lose tail field
- $\mathbf{E} \times \mathbf{B} \rightarrow$  plasma leaves dayside + lost from tail
- (Stronger mantle, diffuse plasmasheet)

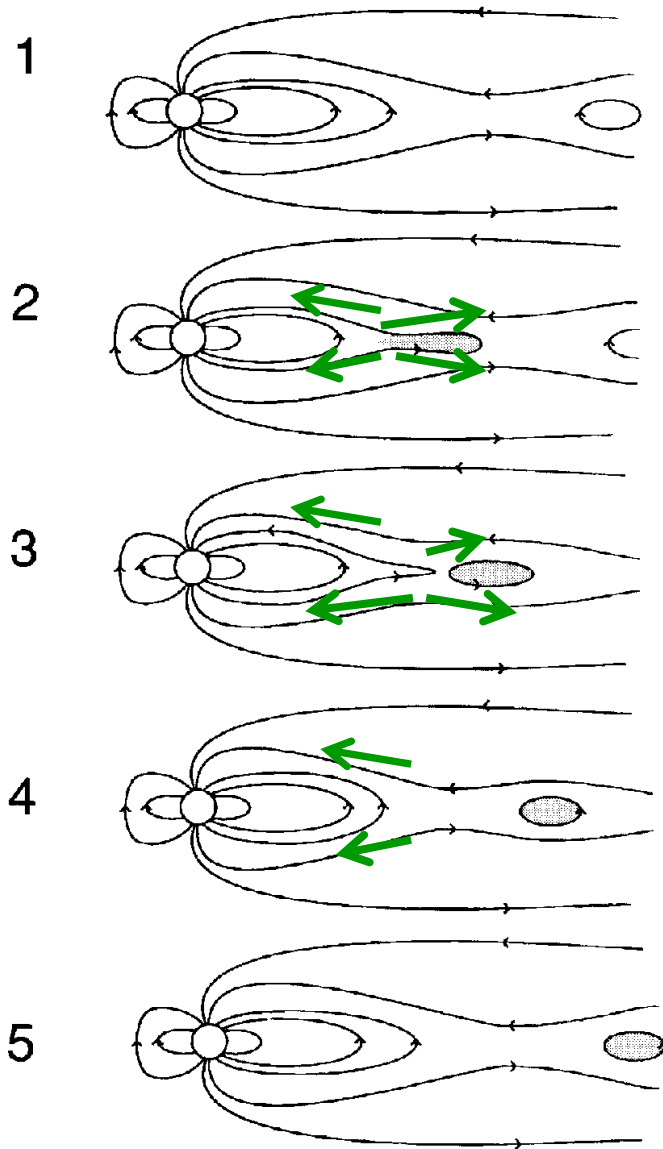
# Southward IMF & dawn-to-dusk $\mathbf{E}_{\text{SW}} \rightarrow$ instability & space weather



- Dayside magnetic reconnection & in tail
- Strip  $\mathbf{B}$  from dayside & build tail field
- $\mathbf{ExB} \rightarrow$  plasma &  $\mathbf{B}$  moves toward neutral sheet
- $\rightarrow$  new reconnection in tail  $\rightarrow$  hot fast plasma to ring current + auroral region &  $\mathbf{B}$  dipolarizes
- plasmoid leaves magnetotail

# Reconnection & Plasmoid Ejection

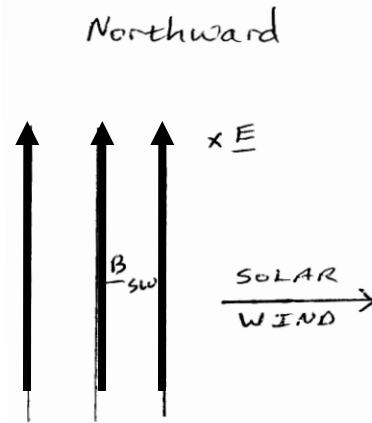
[Hones, 1984]



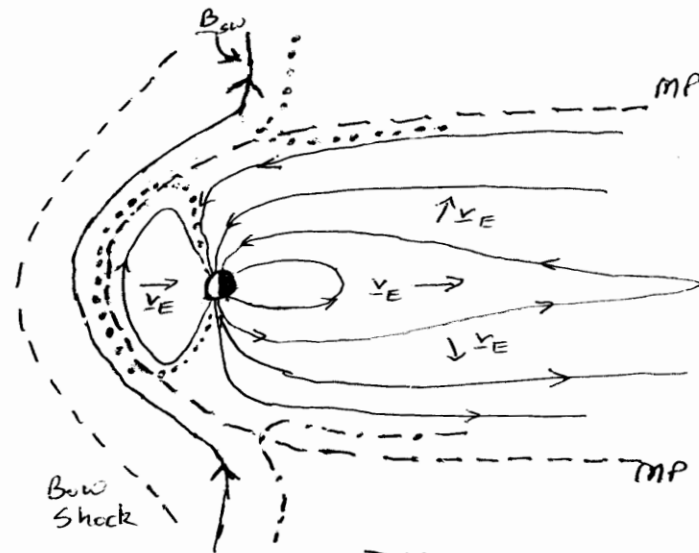
Reconnection outflow of  
hot, fast plasma &  
energetic particles

→ Enhanced ring current,  
aurora,  
increased B variability,  
power line issues,  
increased ionospheric  
ionization

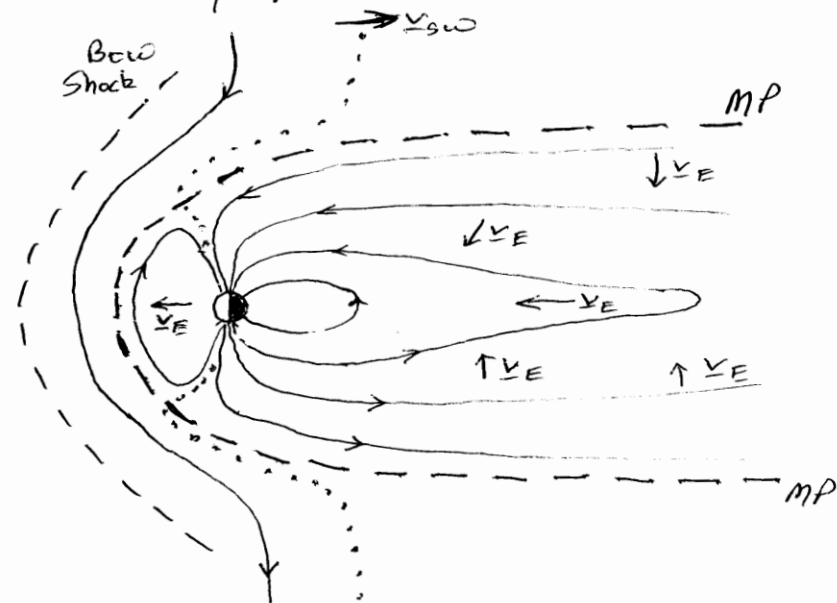
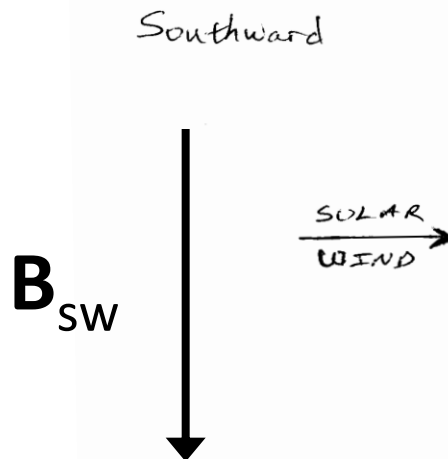
$\times \mathbf{E}_{\text{SW}}$



North  $B_z \rightarrow$  loading phase



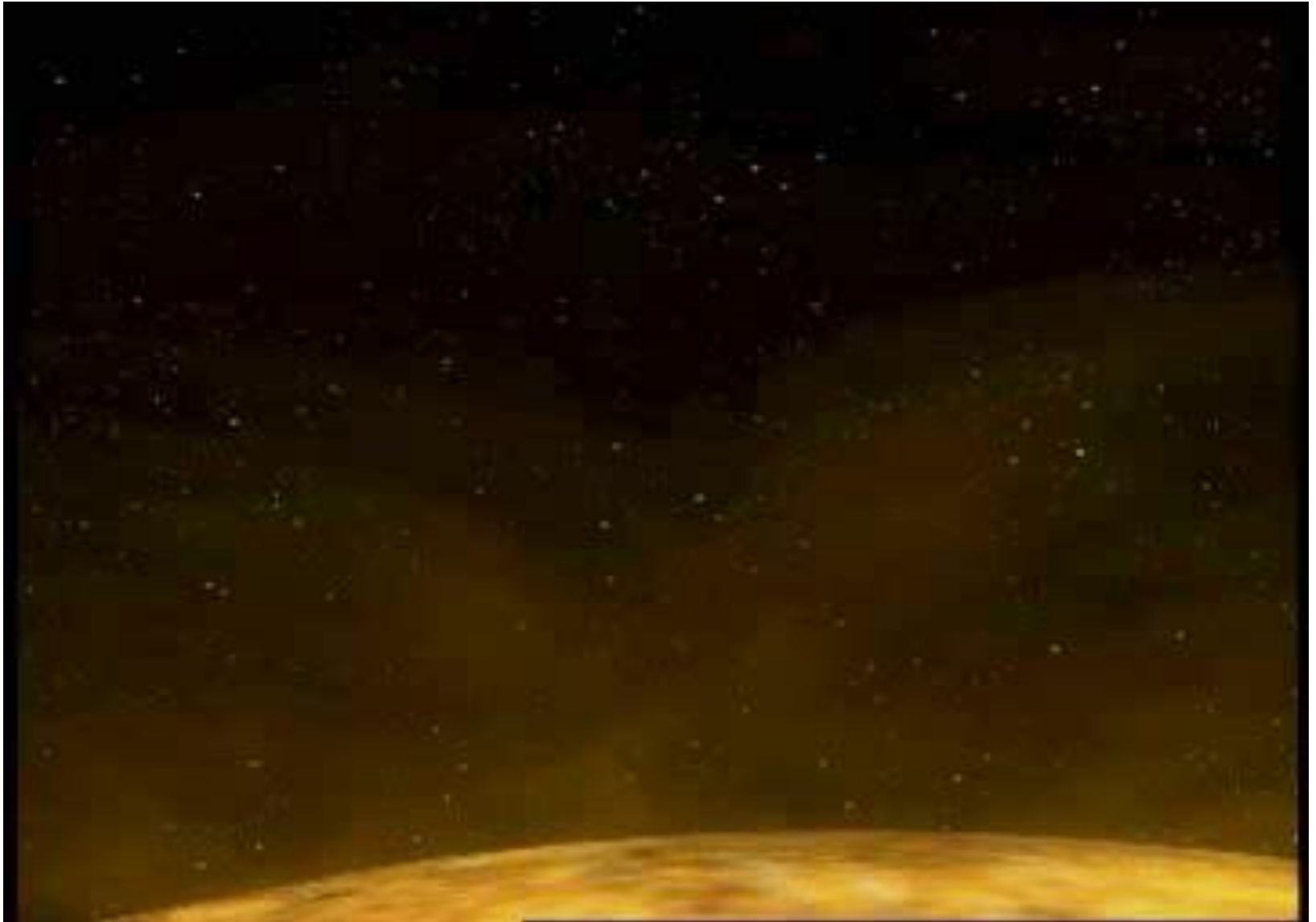
$\cdot \mathbf{E}_{\text{SW}}$



South  $B_z \rightarrow$  instability / unloading with Space Weather events



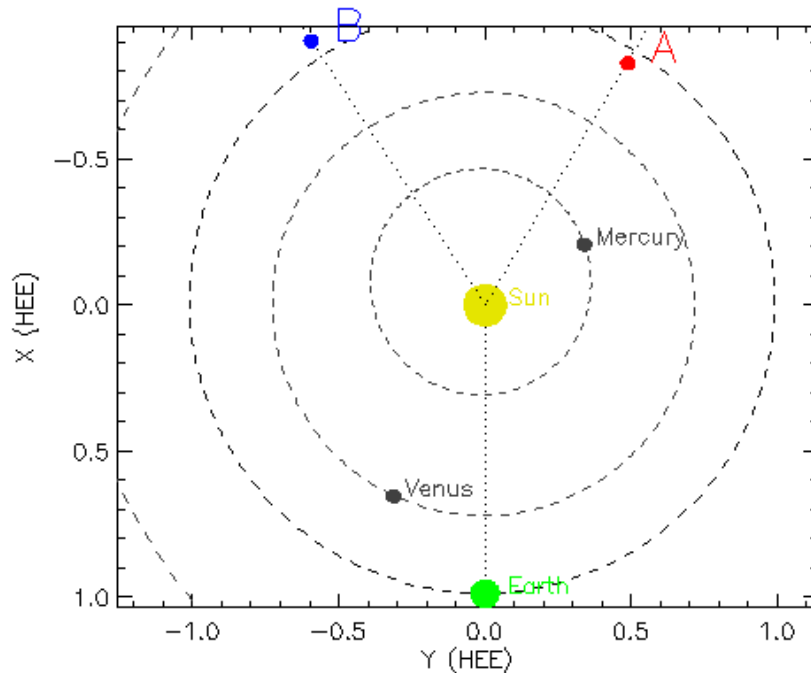
Space weather at Earth mostly due to CMEs & southward Bz  
..... due to ``magnetic reconnection’’



## 4. Prediction of Space Weather

- Need to:
  - Predict motion and properties of CMEs from Sun to Earth.
  - Predict timing and properties of CIRs
  - Predict resulting magnetic reconnection and changes in Earth's magnetosphere and ionosphere, and so the space weather events.
- Next ?

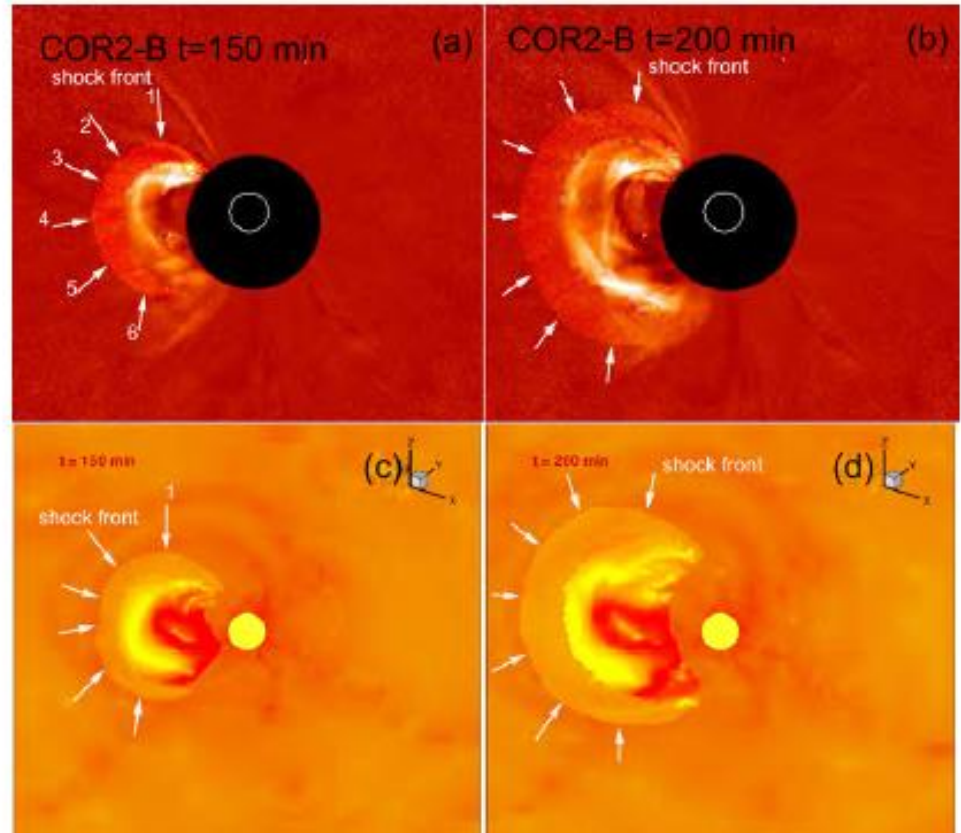
## 5. Sun to Earth: STEREO's 1<sup>st</sup> in situ Interplanetary Type II Burst



- 29 Nov – 1 Dec 2013
- STEREO A and B (double test since  $\Delta r \approx 1$  AU):
  - CME,
  - remote radio,
  - shock crossing at STEREO A only (Langmuir waves, electrons, local radio)
- Excellent agreement from high corona to 1 AU
- $B_z$  prediction .... → space weather prediction

# 5.1 White light predictions from simulations

Data



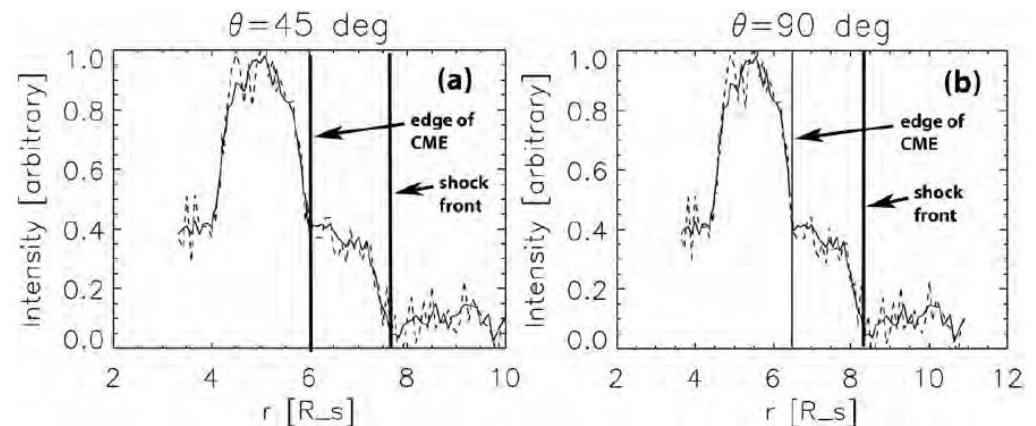
29 Nov 2013  
CME near 20 UT

Prediction  
with  
BATS-R-US

- Predict observable shock
- White light observations &
- predictions agree well.

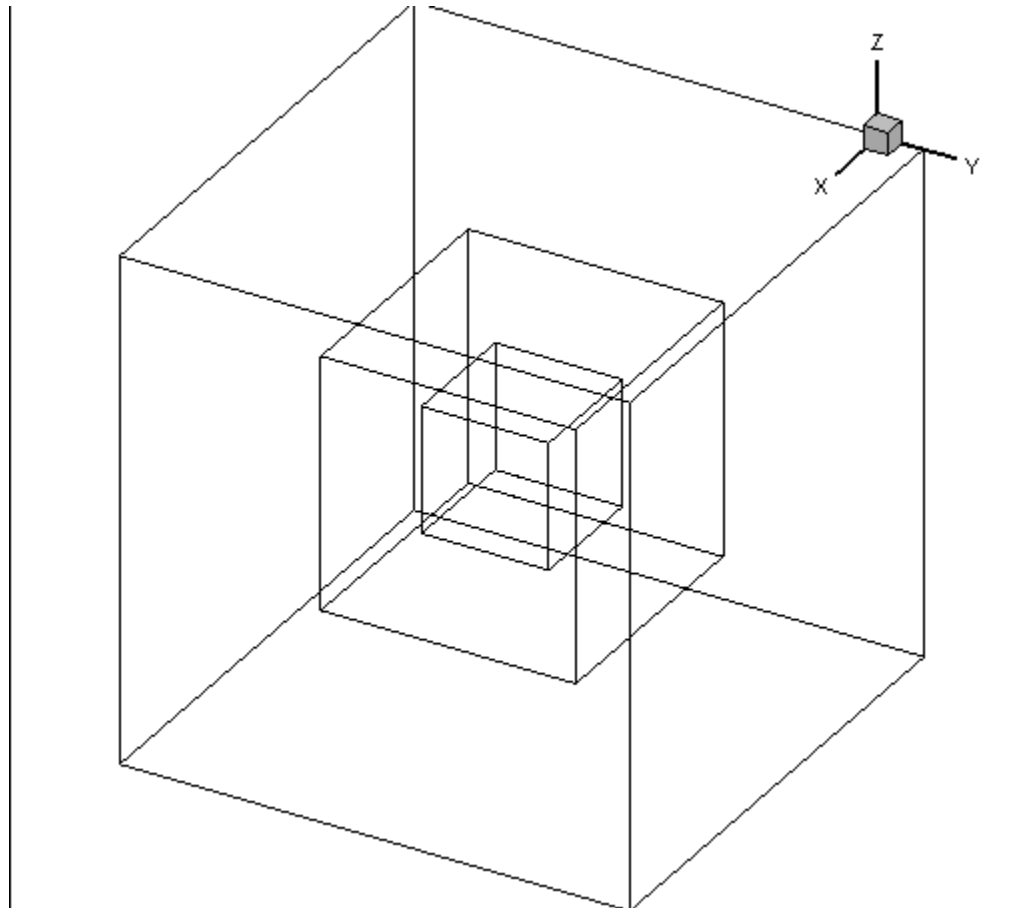
Intensity along rays

➔ Clearly abrupt shocks

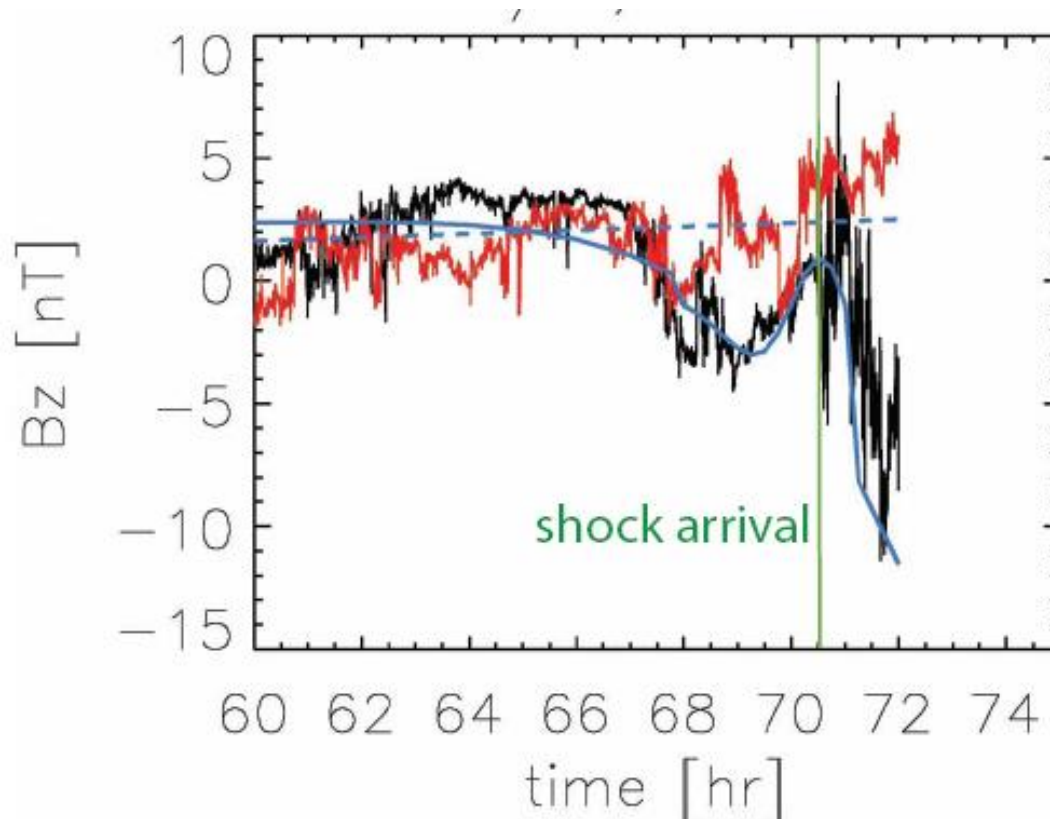


[Schmidt et al., JGR, 2016]

# CME evolution from Sun to STEREO A (29 Nov – 1 Dec 2013)



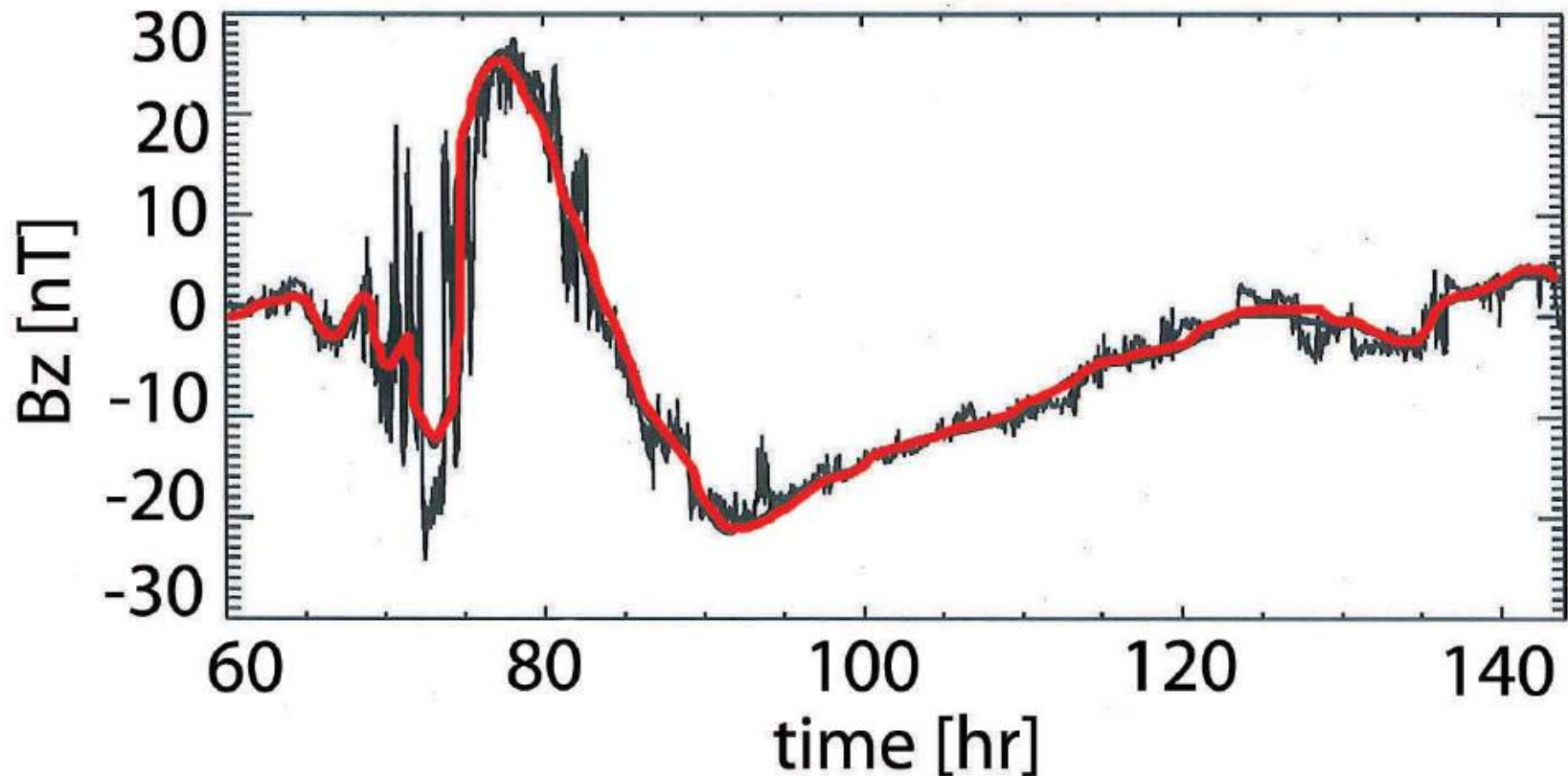
## 5.2 Good $B_z$ and CME predictions at STEREO A



[ Cairns and  
Schmidt, 2015]

- STEREO A data (black) and prediction (blue solid curve)
- ➔ good prediction of  $B_z$  and shock / CME arrival
- ➔ vision intact to predict space weather using type II ...
- STEREO B less good: data (red) and prediction (---)

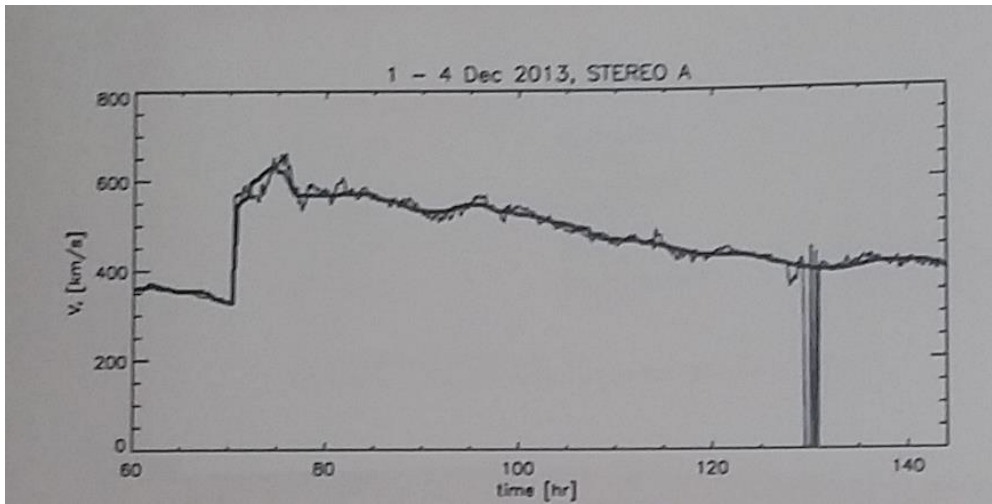
1- 4 Dec. 2013, STEREO A



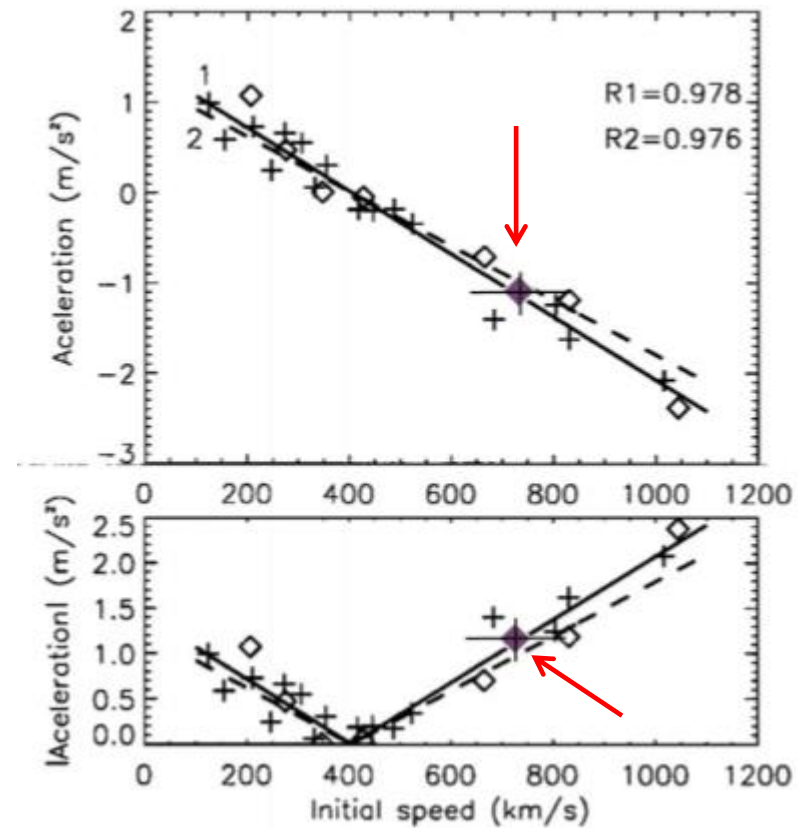
- STEREO A data (black) and prediction (red)
- ➔ very good prediction of  $B_z$  before shock ➔ into CME



# Shock - CME motion & arrival predicted well



- STEREO A data (thin black) and prediction (thick black)



[M. Astore, 2016]

- $V_x$  and shock arrival predicted very well.
- Simulated shock decelerates as predicted by Gopalswamy et al. [2000] data & model for CMEs

## 6. Conclusions

- Most large space weather events at Earth are due to CMEs with southwards  $B_z$  & associated magnetic reconnection .
- Arguably close to accurately simulating CMEs from the Sun to Earth, based on Schmidt, Cairns, et al. case studies
  - White light images can be accurately predicted (CME and shock).
  - CME plasma & field variables can be predicted well, including  $B_z$ .
  - CME arrival, speed & deceleration can agree well with data.
- Agreement → strong support that SWMF / BATS-R-US code (3D MHD) can accurately model 3D corona/wind & CME.
- **Vision & space weather relevance:** → use type II burst (not shown) & white light data-theory iterations to confidently predict in transit CME arrival and  $B_z$ .